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Particle Motion and Continua in Relativity

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LAGRANGIAN APPROACH TO SPINNING OR RADIATING PARTICLE HIGHER-ORDER
 EQUATIONS OF MOTION IN SPECIAL RELATIVITY

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We recognize it worthwhile to distinguish among the higher-order equations of motion those admitting a variational reformulation. Considering Dixon's form of Papapetrou equations either for free particle,

$$p' = 0, S' = 2p\wedge u, \quad (1)$$

or introducing an external force by means of $p' = ku'^2u$, and imposing within the restriction $\|u\| = 1$ a supplementary condition $S'u + ku' = 0$ one regains the Lorentz-Dirac equation, $\mu u' = k(u'' + u'^2u)$, in the second case with constant mass $m = pu$, reversing thus the considerations of A.O.Barut. The Lorentz-Dirac equation, however, cannot be reformulated in terms of a variational principle*. In the presence of another supplementary condition, $Su = 0$, the equations (1) can be given the form $\mu u' = Su''$ (M.Mathisson, 1937), $u \wedge S' = 0$ ($u^2 = 1$). In terms of the star operator $*$, the spin four-vector $\|u\|s = \frac{1}{2}*(u \wedge S)$, and complying with the obvious constraint $su = 0$, this can be rewritten as

$$\mu u' = *(u'' \wedge u \wedge S) \quad (u^2 = 1, s' = 0). \quad (2)$$

Renormalizing the mass by means of $m = m_*(1 - (su)^2/s^2u^2)^{3/2}$, and taking m_* to be an arbitrary but fixed constant we state here that the following third-order equation of motion defining the same set of world lines as (2),

$$m_* \frac{u^2 u' - (u' \wedge u)u}{\|s\|^3 \|u\|^3} = \frac{*(u'' \wedge u \wedge S)}{\|s \wedge u\|^3} - 3 \frac{*(u' \wedge u \wedge S)(s \wedge u') \cdot (s \wedge u)}{\|s \wedge u\|^5},$$

is the Euler-Poisson equation of a parameter-invariant variational problem. Proceeding further in eliminating the spin variables from (2) results in the fourth-order equation $u''' = (p^2/s^2)u'$, ($u^2 = 1$), which under the assumption $p^2 > 0$ was suggested from the higher-order lagrangian point of view by F.Riewe without indicating any direct relationship with Papapetrou equations.

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